

all the pixels along the spectral-dispersion dimension would be summed to obtain the value of the cross-correlation (plus background).

Such on-chip cross-correlation could be performed rapidly because the analytical function could be statically programmed into the APS array and the multiplications could be done simultaneously or nearly so. All of the additions could be done simultaneously by means of a single binning instruction. The charge wells of all the pixels could be connected simultaneously, collecting all the charge outputs from multiplication operations into one "superpixel," the single readout value of which would constitute the cross-correlation value for the given analytical function. For an instrument in which the APS rows were aligned along the spectral-dispersion dimension and in which the image of a spectrograph slit was aligned along the pixel columns and spanned multiple pixel rows, it

would be possible to perform simultaneous cross-correlations for multiple target species by applying, to each pixel row, the analytical function corresponding to one of the target species. A separate readout would be needed for each target species.

In the other hardware implementation, cross-correlations would be computed externally to the APS array. The multiplications and additions would be performed in pipeline fashion. If the APS-array outputs were analog, then programmable analog signals representing the analytical functions would be synthesized in phase with the corresponding stream of analog APS-array outputs and the multiplications and additions would be performed by relatively inexpensive, commercially available analog mixing and filtering circuits, respectively. If the APS-array outputs were digital, the cross-correlations could be computed by a digital signal processor. Ordinarily, the analog

approach would be preferable because the analog operations can be performed much more rapidly than can the corresponding digital multiplications and additions.

This work was done by Gregory Bearman, Michael Pelletier, and Suresh Seshadri of Caltech for NASA's Jet Propulsion Laboratory. Further information is contained in a TSP (see page 1).

In accordance with Public Law 96-517, the contractor has elected to retain title to this invention. Inquiries concerning rights for its commercial use should be addressed to:

*Innovative Technology Assets Management
JPL*

*Mail Stop 202-233
4800 Oak Grove Drive
Pasadena, CA 91109-8099
(818) 354-2240*

E-mail: iaoffce@jpl.nasa.gov

Refer to NPO-30912, volume and number of this NASA Tech Briefs issue, and the page number.

Prioritizing Scientific Data for Transmission

NASA's Jet Propulsion Laboratory, Pasadena, California

A software system has been developed for prioritizing newly acquired geological data onboard a planetary rover. The system has been designed to enable efficient use of limited communication resources by transmitting the data likely to have the most scientific value. This software operates onboard a rover by analyzing collected data, identifying potential scientific targets, and then using that information to prioritize data for transmission to Earth. Currently, the system is focused on the analysis of acquired images, although the general techniques are applicable to a wide range of data modalities. Image prioritization is performed using two main

steps. In the first step, the software detects features of interest from each image. In its current application, the system is focused on visual properties of rocks. Thus, rocks are located in each image and rock properties, such as shape, texture, and albedo, are extracted from the identified rocks. In the second step, the features extracted from a group of images are used to prioritize the images using three different methods: (1) identification of key target signature (finding specific rock features the scientist has identified as important), (2) novelty detection (finding rocks we haven't seen before), and (3) representative rock sampling (finding

the most average sample of each rock type). These methods use techniques such as K-means unsupervised clustering and a discrimination-based kernel classifier to rank images based on their interest level.

This program was written by Rebecca Castano, Robert Anderson, Tara Estlin, Dennis DeCoste, Daniel Gaines, Dominic Mazzoni, Forest Fisher, and Michele Judd of Caltech for NASA's Jet Propulsion Laboratory. Further information is contained in a TSP (see page 1).

This software is available for commercial licensing. Please contact Don Hart of the California Institute of Technology at (818) 393-3425. Refer to NPO-40265.

Determining Sizes of Particles in a Flow From DPIV Data

The same equipment would be used to measure sizes as well as velocities.

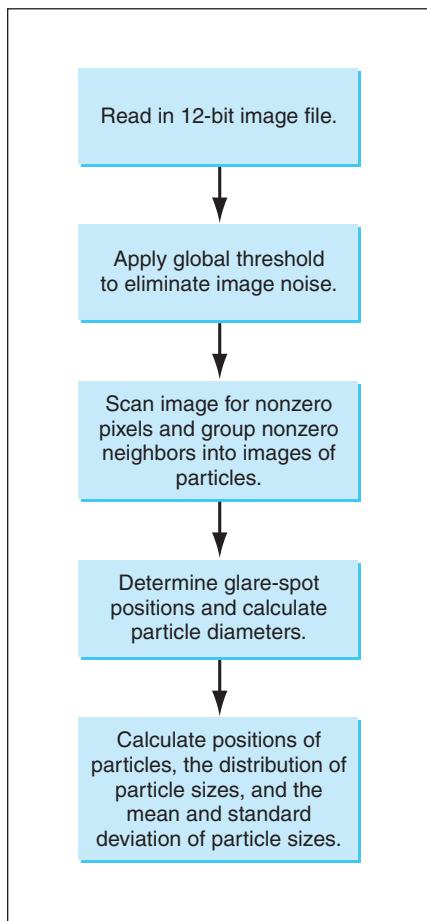
John H. Glenn Research Center, Cleveland, Ohio

A proposed method of measuring the size of particles entrained in a flow of a liquid or gas would involve utilization of data from digital particle-image velocimetry (DPIV) of the flow. That is to say, with proper design and operation of

a DPIV system, the DPIV data could be processed according to the proposed method to obtain particle sizes in addition to particle velocities. As an additional benefit, one could then compute the mass flux of the entrained particles

from the particle sizes and velocities.

As in DPIV as practiced heretofore, a pulsed laser beam would be formed into a thin sheet to illuminate a plane of interest in a flow field and the illuminated plane would be observed by means of a



The **Algorithm of the Proposed Method** would process DPIV data to extract information on particle sizes. The size information would be in addition to the velocity information extracted from the data by use of previously developed DPIV software.

charge-coupled device (CCD) camera aimed along a line perpendicular to the illuminated plane. Unlike in DPIV as practiced heretofore, care would be taken to polarize the laser beam so that its electric field would lie in the illuminated plane, for the reason explained in the next paragraph.

The proposed method applies, more specifically, to transparent or semitransparent spherical particles that have an index of refraction different from that of the fluid in which they are entrained. The method is based on the established Mie theory, which describes the scattering of light by diffraction, refraction, and specular reflection of light by such particles. In the case of a particle illuminated by polarized light and observed in the arrangement described in the preceding paragraph, the Mie theory shows that the image of the particle on the focal plane of the CCD camera includes two glare spots: one attributable to light reflected toward the camera and one attributable to light refracted toward the camera. The distance between the glare spots is a known function of the size of the particle, the indices of refraction of the particle material, and design parameters of the camera optics. Hence, the size of a particle can be determined from the distance between the glare spots.

The proposed method would be implemented in an algorithm that would automatically identify, and measure the distance between, the glare spots for each particle for which a suitable image has been captured in a DPIV image frame. The algorithm (see figure) would begin with thresholding of data from the entire image frame to reduce noise, thereby facilitating discrimination of particle images from the background and aiding in the separation of overlapping particles. It is important not to pick a threshold level so high that the light intensity between a given pair of glare spots does not fall below the threshold value, leaving the glare spots disconnected.

The image would then be scanned in a sequence of rows and columns of pixels to identify groups of adjacent pixels that contain nonzero brightnesses and that

are surrounded by pixels of zero brightness. Each such group would be assumed to constitute the image of one particle. Each such group would be further analyzed to determine whether the image was saturated; saturated particle images must be rejected because the locations of glare spots in saturated images cannot accurately be determined. Within each unsaturated particle image, the centroids (deemed to be the locations) of the glare spots would be determined by means of gradients of brightness distributions and three-point horizontal and three-point vertical Gaussian estimates based on the brightness values of the brightest pixels and the pixels adjacent to them. If the brightness of a given particle image contained only one peak, then it would be assumed that a second glare spot did not exist and that image would be rejected.

Once the centroids had been estimated for all particle images for which it was possible to do so, the positions of the particles and the distances between their centroids would be computed. As described above, the size of each particle would then be computed from the distance between its centroids. Finally, the distribution, mean, and standard deviation of sizes would be computed for the collection of particle images that survived to the final stage of the centroid-estimation process.

This work was done by M. P. Wernet of Glenn Research Center and A. Mielke and J. R. Kadambi of Case Western Reserve University. Further information is contained in a TSP (see page 1).

Inquiries concerning rights for the commercial use of this invention should be addressed to NASA Glenn Research Center, Commercial Technology Office, Attn: Steve Fedor, Mail Stop 4-8, 21000 Brookpark Road, Cleveland, Ohio 44135. Refer to LEW-17340.

Faster Processing for Inverting GPS Occultation Data

NASA's Jet Propulsion Laboratory, Pasadena, California

A document outlines a computational method that can be incorporated into two prior methods used to invert Global Positioning System (GPS) occultation data [signal data acquired by a low-Earth-orbiting satellite as either this or the GPS satellite rises above or falls below the horizon] to obtain information on altitude-depen-

dent properties of the atmosphere. The two prior inversion methods, known as back propagation and canonical transform, are computationally expensive because for each occultation, they involve numerical evaluation of a large number of diffraction-like spatial integrals. The present method involves an angular-spectrum-based phase-ex-

trapolation approximation in which each data point is associated with a plane-wave component that propagates in a unique direction from the orbit of the receiving satellite to intersect a straight line tangent to the orbit at a nearby point. This approximation enables the use of fast Fourier transforms (FFTs), which apply only to data